Eastwood High School

Name:

Unit 1 Our Dynamic Universe



Summary Notes

1.1 Equations of Motion

Vectors

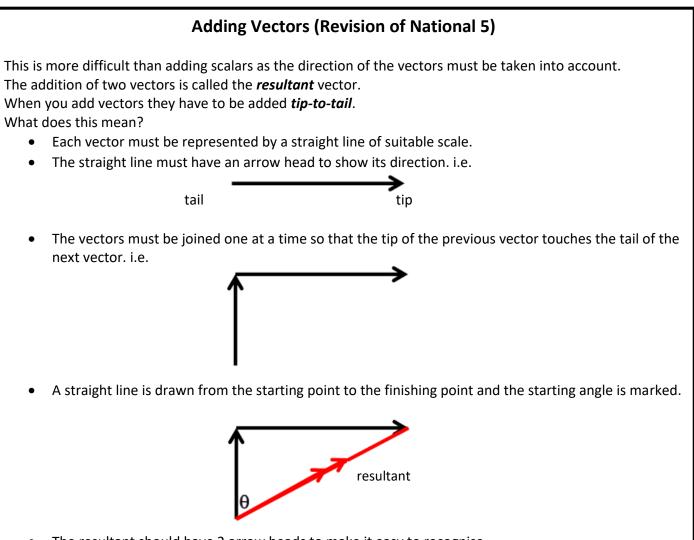
Vectors and Scalars (Revision of National 5)

It is possible to split up quantities in physics into two distinct groups, those that need a direction and those that don't. Some are obvious - it makes sense that force has direction; you can push or pull but you need to specify the direction.

It would be nonsense to give a direction to mass. To say: "Its mass was 5kg East" just isn't right. It is important that you are familiar with which quantity falls into which grouping.

A scalar is a quantity that can be described by just a **magnitude** and a unit. e.g. time - 30 s, mass - 20 kg.

A vector is a quantity that is fully described with a **magnitude and** *direction*. e.g. force : 50 N downwards velocity : 20 ms⁻¹ East.



- The resultant should have 2 arrow heads to make it easy to recognise.
- If using a scale diagram the length and direction of this straight line gives the resultant vector.
- Alternatively you can use trigonometry and SOHCAHTOA or the sine or cosine rule to calculate the resultant.

Distance and Displacement (Revision of National 5)

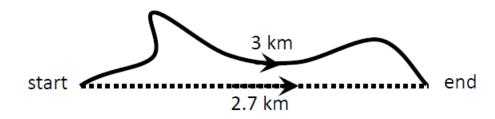
The **distance** travelled by an object is the **sum of the distances** of each stage of the journey.

Since each stage has a different direction, the total distance has no single direction and therefore distance is a scalar.

The **displacement** of an object is the **shortest route between the start and finish point measured in a straight line**.

Displacement has a direction and is a vector.

Consider the journey below. A person walks along a path (solid line) from start to end.



They will have walked further following the path than if they had been able to walk directly from start to end in a straight line (dashed line).

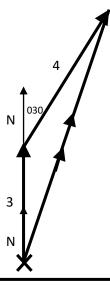
The solid line denotes the **distance** = 3km. The dashed line denotes the **displacement** = 2.7 km East

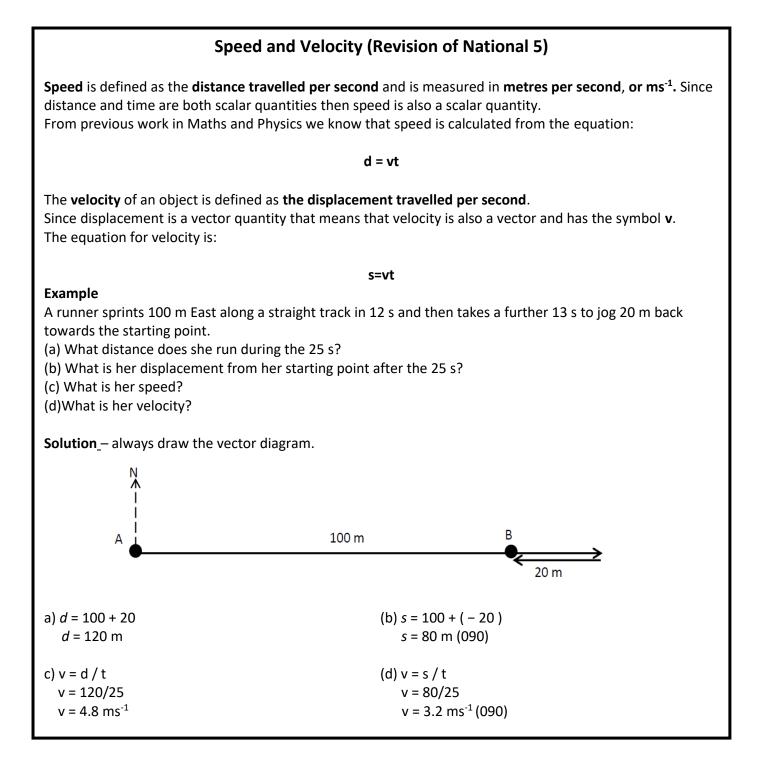
Example

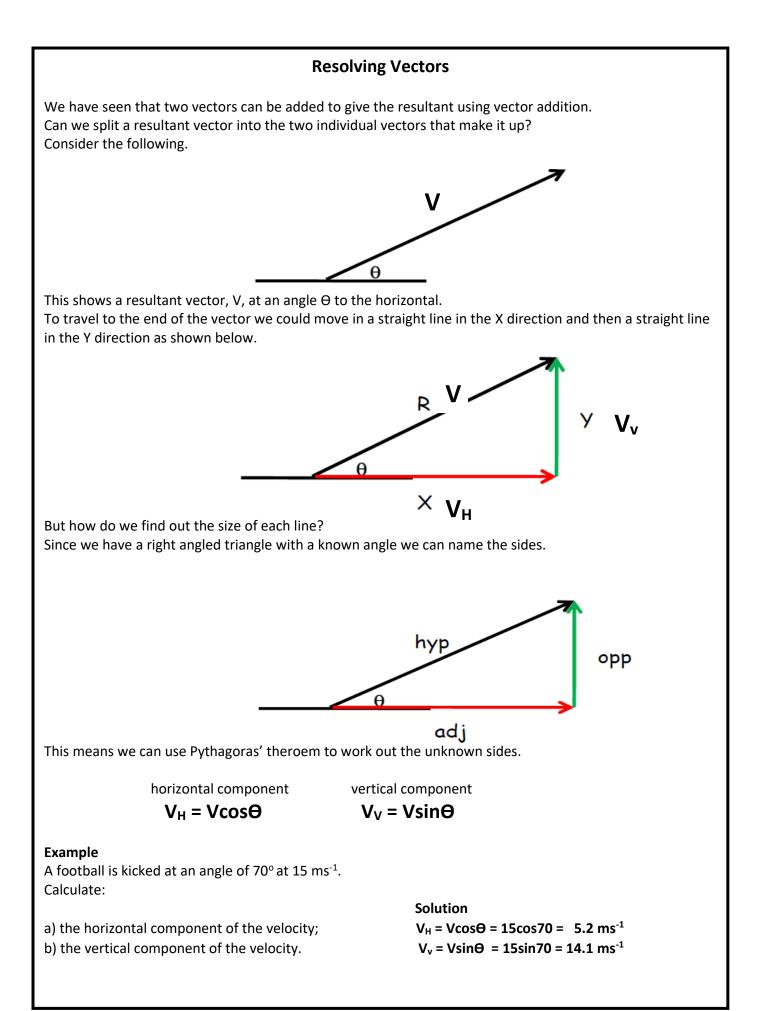
A woman walks her dog 3 km due North (000) and then 4 km (030). Find her a) distance travelled b) displacement.

Solution – Use a ruler to measure the lengths of the vectors and a protractor to measure the bearing.

- Choose an appropriate scale e.g. 1cm : 1km
- Mark the start point with an X, draw a North line and draw the first vector.
- Draw a North line at the tip of this vector and now draw the second vector (tip to tail)
- Draw the resultant vector from start to end using the double arrow.
- Measure the length of the line and the bearing.
 When measuring bearings remember from START CLOCKWISE from NORTH







The Equations of Motion (s u v a t)

The equations of motion can be applied to any object moving with constant acceleration in a straight line. You must be able to:

- select the correct formula;
- identify the symbols and units used;
- carry out calculations to solve problems of real life motion; and
- carry out experiments to verify the equations of motion.

$$v = u + at$$
 $s = ut + \frac{1}{2} at^2$ $v^2 = u^2 + 2as$

where

- s = displacement
- u = initial velocity
- v = final velocity
- t = time

a = acceleration

Equation of Motion 1 :

v = u + at

Example

A racing car starts from rest and accelerates uniformly in a straight line at 12 ms⁻² for 5.0 s. Calculate the final velocity of the car.

Solution LIST suvat

s = ? m u = 0 ms⁻¹ (rest) v = ? ms⁻¹ v = u + at a = 12 ms⁻² v = 0 + (12 x5.0) t = 5.0 s v = 0 + 60 v = 60 ms⁻¹



The 3 Equations of Motion (ctd) **Equation of Motion 2:** $s = ut + \frac{1}{2} at^2$ A speedboat travels 400 m in a straight line when it accelerates uniformly from 2.5 ms⁻¹ in 10 s. Calculate the acceleration of the speedboat. $s = ut + \frac{1}{2} at^2$ u = 2.5 ms⁻¹ $400 = (2.5 \times 10) + (0.5 \times a \times 10^2)$ 400 = 25 + 50a 50a = 400 - 25 = 375 a = 375/50 a = 7.5 ms⁻² **Equation of Motion 3:** $v^2 = u^2 + 2as$



Example

Example

Solution

 $v = ? ms^{-1}$

 $a = ? ms^{-2}$

s = 400 m

t = 10 s

A rocket is travelling through outer space with uniform velocity. It then accelerates at 2.5 ms⁻² in a straight line in the original direction, reaching 100 ms⁻¹ after travelling 1875 m. Calculate the rocket's initial velocity.

Solution

u = ? ms ⁻¹	$v^2 = u^2 + 2as$
v = 100 ms ⁻¹	100 ² = u ² + (2 x 2.5 x 1875)
a = 2.5 ms ⁻²	10 000 = u2 +9375
s = 1875 m	u ² = 10 000 - 9375
t = ? s	u = 25ms ⁻¹
	u = 25 ms ⁻¹

The 3 Equations of Motion with Decelerating Objects

When an object decelerates its velocity decreases. If the vector quantities in the equations of motion are positive, we represent the decreasing velocity by use of a negative sign in front of the acceleration value.

Example 1

A car, travelling in a straight line, **decelerates** uniformly at 2.0 ms⁻² from 25 ms⁻¹ for 3.0 s. Calculate the car's velocity after the 3.0 s.

Solution

u = 25 ms⁻¹ v = ? a = -2.0 ms⁻² s = ?m t = 3.0 s v = u + at v = 25 + (-**2.0** x 3.0) v = 25 + (-6.0) v = 19 ms⁻¹



Example 2

A greyhound is running at 6.0 ms⁻¹. It **decelerates** uniformly in a straight line at 0.5 ms⁻² for 4.0 s. Calculate the displacement of the greyhound while it was decelerating.

Solution

u = 6.0 ms ⁻¹	$s = ut + \frac{1}{2} at^{2}$
v = ? ms ⁻¹	$s = (6.0 \times 4.0) + (0.5 \times -0.5 \times 4.0^2)$
a = -0.5 ms ⁻²	s = 24 + (-4.0)
s = ?m	s = 20 m
t = 4.0 s	



Example 3

A curling stone leaves a player's hand at 5.0 ms⁻¹ and **decelerates** uniformly at 0.75ms⁻² in a straight line for 16.5 m until it strikes another stationary stone.

Calculate the velocity of the decelerating curling stone at the instant it strikes the stationary one.

Solution

u = 5.0 ms ⁻¹	$v^2 = u^2 + 2as$
v = ? ms ⁻¹	v ² = 5.0 ² + (2 x -0.75 x 16.5)
a = -0.75 ms ⁻²	v ² = 25 + (-24.75)
s = 16.5 m	v = √0.25
t = ?s	v = 0.5 ms ⁻¹



Graphing Motion

Graphs

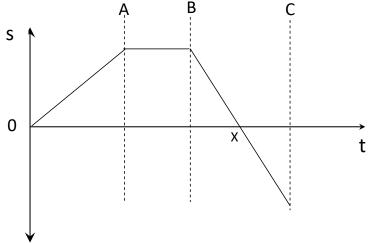
In all areas of science, graphs are used to display information. Graphs are an excellent way of giving information, especially to show relationships between quantities. In this section we will be examining three types of motion-time graphs.

> Displacement-time graphs Velocity-time graphs Acceleration-time graphs

If you have an example of one of these types of graph then it is possible to draw a corresponding graph for the other two factors.

Displacement – time graphs

This graph represents how far an object is from its starting point at some known time. Because displacement is a vector it can have positive and negative values. (+ve and –ve will be opposite directions from the starting point).



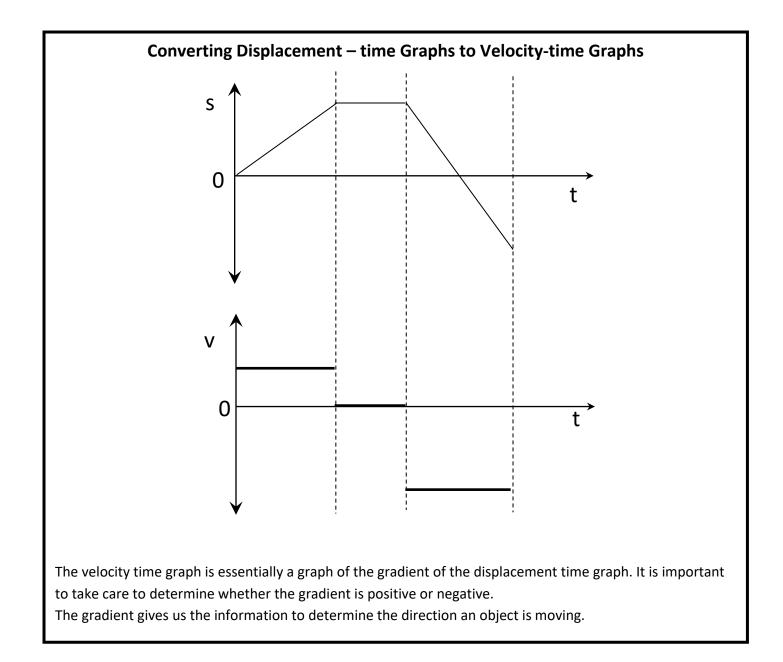
OA – the object is moving away from the starting point. It is moving a constant displacement each second. This is shown by the constant gradient. What does this mean?

gradient =
$$\frac{\text{displacement}}{\text{time}}$$
 = velocity

We can determine the velocity from the gradient of a displacement time graph.

AB – the object has a constant displacement so is not changing its position, therefore it must be at rest. The gradient in this case is zero, which means the object has a velocity of zero [at rest]

BC – the object is now moving back towards the starting point, reaching it at time x. It then continues to move away from the start, but in the opposite direction. The gradient of the line is negative, indicating the change in direction of motion.



Velocity – time Graphs

It is possible to produce a velocity time graph to describe the motion of an object. All velocity time graphs that you encounter in this course will be of objects that have constant acceleration.

Scenario: The Bouncing Ball

Lydia fires a ball vertically into the air from the ground. The ball reaches its maximum height, falls, bounces and then rises to a new, lower, maximum height.

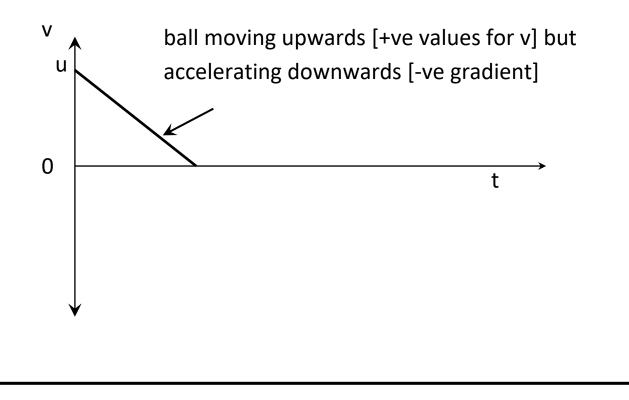
What will the velocity time graph for this motion look like?

First decision: The original direction of motion is up so upwards is the positive direction

Part One of Graph

Now we need to think, what is happening to the velocity?

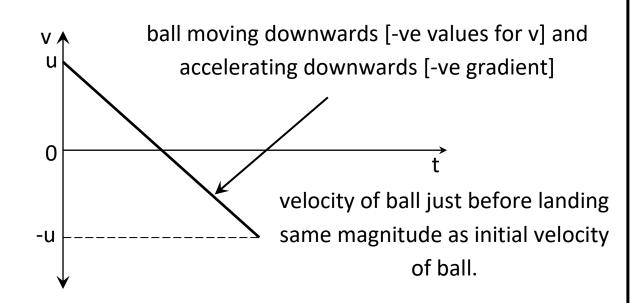
The ball will be slowing down whilst it is moving upwards, having a velocity of zero when it reaches maximum height. The acceleration of the ball will be constant if we ignore air resistance.



Velocity – time Graphs (continued)

Part Two of Graph

Once the ball reaches its maximum height it will begin to fall downwards. It will accelerate at the same rate as when it was going up. The velocity of the ball just before it hits the ground will be the same magnitude as its initial velocity upwards

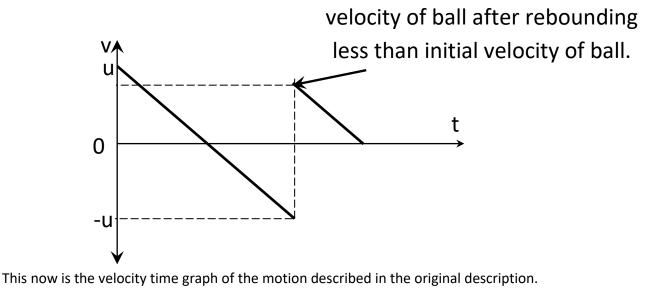


Part Three of Graph

The ball has now hit the ground. At this point it will rebound and begin its movement upwards.

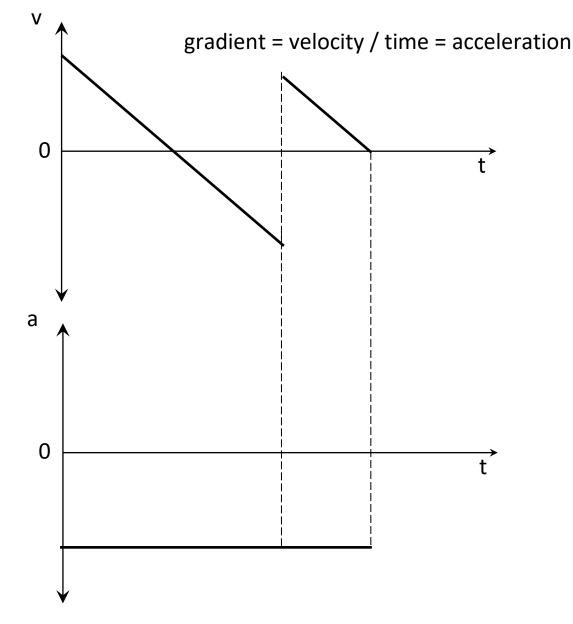
In reality there will be a finite time of contact with the ground when the ball compresses and regains its shape. In this interpretation we will regard this time of contact as zero. This will result in a disjointed graph.

The acceleration of the ball after rebounding will be the same as the initial acceleration. The two lines will be parallel.



Converting Velocity – time Graphs to Acceleration – time Graphs

What is important in this conversion is to consider the gradient of the velocity-time graph line. In our example the gradient of the line is constant and has a **negative** value. This means for the entire time sampled the acceleration will have a single **negative** value.



All acceleration time graphs you are asked to draw will consist of horizontal lines, either above, below or on the time axis.

Reminder from National 5

The area under a speed time graph is equal to the distance travelled by the object that makes the speed time graph.

In this course we are dealing with vectors so the statement above has to be changed to:

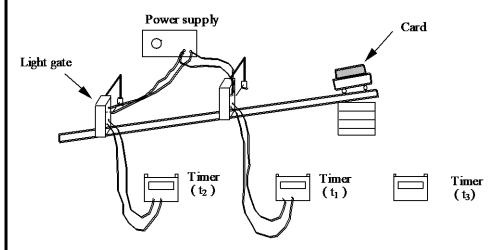
The area under a **velocity** time graph is equal to the **displacement of** the object that makes the speed time graph.

Any calculated areas that are below the time axis represent **negative** displacements.

Experiment : Measuring Acceleration Down a Slope

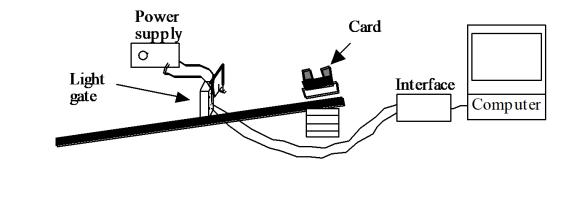
Aim: To calculate the acceleration of a trolley using 2 light gates.

Apparatus: Slope, trolley with mask, 2 light gates and electronic timers.



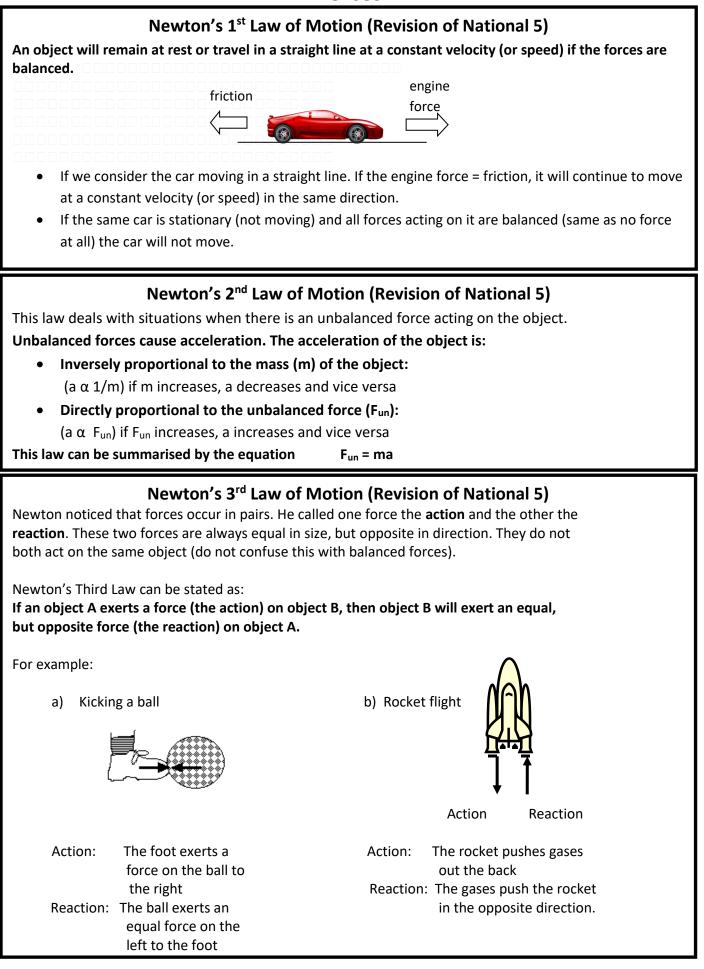
- Release the trolley from the top of the slope
- Start timer 3 once the trolley cuts the first light gate and stop it once it cuts the second light gate
- Using the length of the card and values from t₁ and t₂ calculate the initial (u) and final velocity (v)
- Using v, u and t₃ calculate the acceleration of the trolley
- Repeat this experiment 5 times, calculate a mean acceleration and associated random uncertainty.

This experiment can also be done with a 'double mask' to measure the acceleration at a single point with a single light gate.



1.2 Forces, Energy and Power

Forces



Resultant Forces – Horizontal

When several forces act on one object, they can be replaced by one force which has the same effect. This single force is called the resultant or unbalanced force. Remember that Friction is a resistive force which acts in the **opposite** direction to motion. By increasing friction you reduce the unbalanced force which in turn reduced the acceleration.

200N

100N

1200N

Example: Horizontal

A motorcycle and rider of combined mass 650 kg provide an engine force of 1200 N. The friction between the road and motorcycle is 100N and the drag value = 200N.

Calculate:

- a) the unbalanced force acting on the motorcycle
- b) the acceleration of the motorcycle

Solution

a) Draw a free body diagram

F_{un} = 1200 - 200 - 100 Fun = 900 N This 900 N force is the resultant of the 3 forces b) F_{un} = 900 N $F_{un} = ma$ a = ? 900 = 650 x a

 $a = 1.38 \text{ ms}^{-2}$ m = 650 kg

Resultant Forces – Vertical (Rocket)

Example

At launch, a rocket of mass 20 000 kg accelerates off the ground at 12 ms⁻² (ignore air resistance)

- a) Use Newton's 3rd law of motion to explain how the rocket gets off the ground.
- b) Draw a free body diagram to show all the vertical forces acting on the rocket as it accelerates upwards.
- c) Calculate the engine thrust of the rocket which causes the acceleration of 12ms⁻².

Solutions

- a) The rocket pushes the gas out the back downwards (action) and the gas pushes the rocket upwards (reaction).
- b) Engine thrust Weight Engine thrust = ? c) Calculate Fun and W $F_{un} = m a$ F_{un} = 20 000 x 12 F_{un} = 240 000 N F_{un} = ma W = mgm = 20000 kg W = 20 000 x 9.8 W = 196 000 N Fun = upward force (thrust) – downwards force (Weight) 240 000 = thrust - 196 000 W = mg <u>thrust = 436 000 N</u>

Resultant Forces – Vertical (Lift)

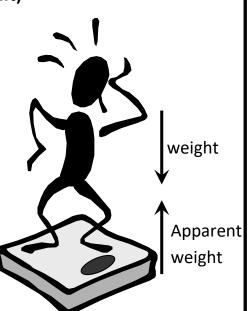
Have you noticed that when you are in a lift you experience a strange feeling when the lift starts to move and as it begins to slow to a stop. However when the lift is in the middle of its journey you cannot tell if you are moving at all.

This is because at the start and end of the journey you will experience an acceleration and consequently an unbalanced force. This unbalanced force is what you 'feel'.

When you stand on a set of scales (Newton Balance) the reading on the scales is actually measuring the **upwards** force. This is the force the scales exert on you.

We will call this the Apparent Weight.

Now this is fine when you are in your bathroom trying to find your weight.



Normally, you and your bathroom scales will be stationary and so your weight will be equal to the upwards force (balanced forces).

When you weigh yourself when you are accelerating the reading on the scales will <u>not</u> be your weight. The reading will give you an indication of the unbalanced force acting on you, which could then be used to calculate an acceleration. This unbalanced force could be acting up or down depending on the magnitude and direction of the acceleration.

The value of the Apparent Weight will be equal to the Tension T in the cable of the lift.

The table below explains the motion of a lift as it moves from an upper floor to a lower floor and then back to the original floor.

Motion	Comparing W and T	Unbalanced Force F	Sensation in lift
Stationary at top	W=T	F = 0	
Accelerating down	T decreases W > T	F = W - T	Lighter
Constant velocity down	T increases W = T	F = 0	
Decelerating down	T increases T > W	F = T – W	Heavier
Stationary at bottom	T decreases T = W	F = 0	
Accelerating up	T increases T > W	F = T – W	Heavier
Constant speed up	T decreases T = W	F = 0	
Decelerating up	T decrease W > T	F = W - T	Lighter

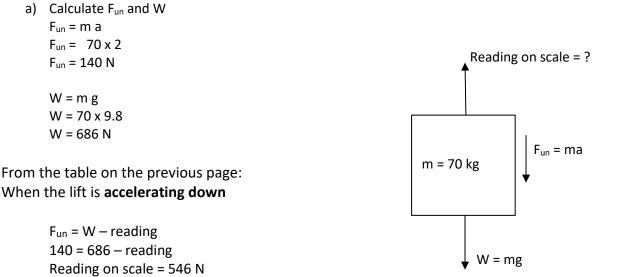
Resultant Forces – Vertical (Lift continued)

Example

A man of mass 70 kg stands on a set of bathroom scales in a lift. Calculate the reading on the scales when the lift is accelerating downwards at 2 ms⁻².

Solution

Remember that the reading on the scales = apparent weight = tension in the cable



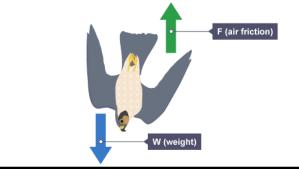
Terminal Velocity

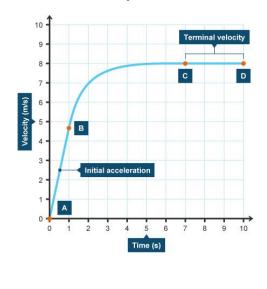
Newton's First Law applies particularly well under freefall conditions.

Sky divers who jump from a plane do not continue to accelerate to the ground at ever increasing speeds – if they did, how could they create spectacular formations? After jumping out of a plane the skydiver begins to accelerate downwards at 9.8ms⁻². However, as the skydiver's speed increases the frictional force due to air resistance also increases. There comes a point when the skydiver's upward frictional force equals the skydiver's downward weight. The forces are now balanced. If the vertical forces on the skydiver are balanced then according to Newton's First Law, the skydiver will continue downwards at a constant speed. This constant speed (or velocity) is known as the **terminal velocity**.

The bird in the picture below has the same issue.

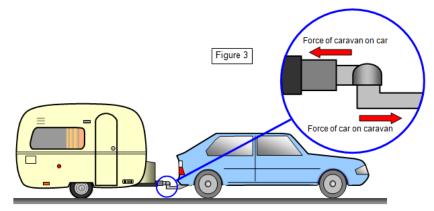
The bird will fall on its prey with a constant velocity due to balanced forces. The graph shows how the velocity of the bird would change as it approached terminal velocity.





Internal Forces: Tension

An example of an internal force is the tension in the towbar (magnified below) when a car is pulling a caravan.



In higher physics, a common question in the SQA exam you are asked is to calculate the tension between the two objects.

Example

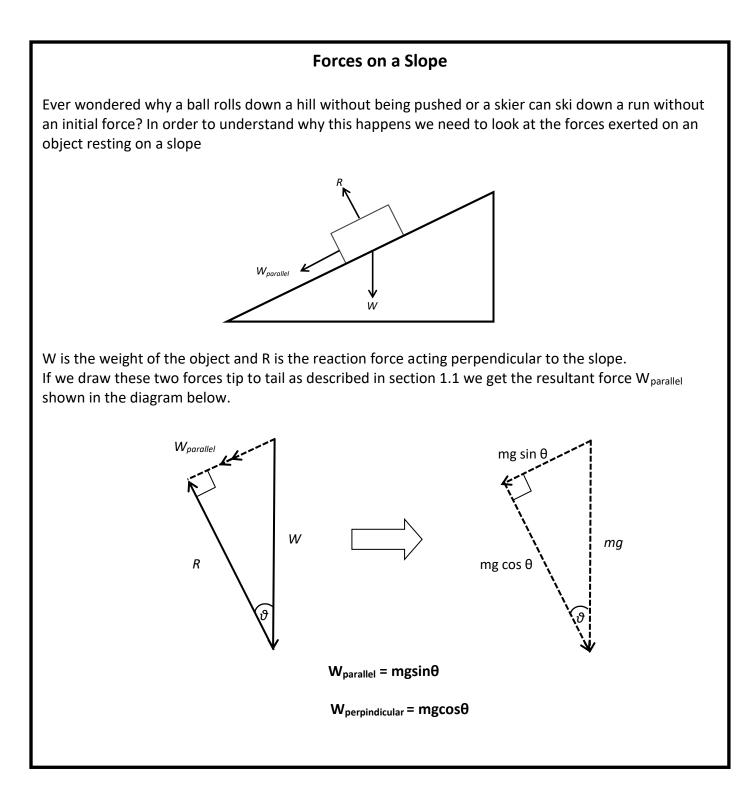
A car of mass 700 kg pulls a 500 kg caravan with a constant engine thrust of 3.6 kN. Calculate the tension in the towbar during the journey (ignoring friction)

Solution: HINT Calculate the acceleration of the whole system using Fun = ma

F _{un} = 3600 N	a = F _{un} / m
m = 500 + 700 = 1200 kg	a = 3600 / 1200
a = ?	a = 3 ms ⁻²

Use this acceleration to calculate the tension and use the mass of the caravan only as this is the mass of the object being pulled.

T = ?	T = ma
m = 500 kg	T = 500 x 3
a = 3ms ⁻²	T = 1500 N

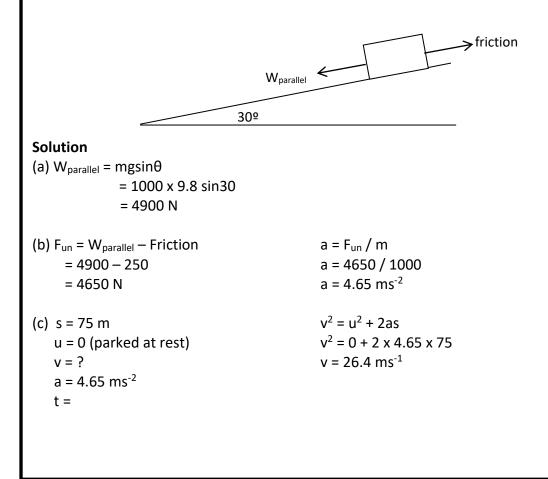


Forces on a Slope (continued)

Example

A car of mass 1000 kg is parked on a hill. The slope of the hill is 20° to the horizontal. The brakes on the car fail. The car runs down the hill for a distance of 75 m until it crashes into a hedge. The average force of friction on the car as it runs down the hill is 250 N.

- (a) Calculate the component of the weight acting down (parallel to) the slope.
- (b) Find the acceleration of the car.
- (c) Calculate the speed of the car just before it hits the hedge.



Energy

Conservation of Energy

One of the fundamental principles of Physics is that of conservation of energy.

Energy cannot be created or destroyed, only converted from one form to another.

Work is done when converting from one form of energy to another. Power is a measure of the rate at which the energy is converted.

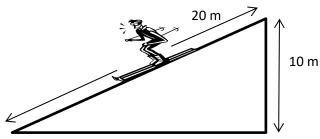
There are a number of equations for the different forms of energy:

Ew	=	Fd
Eĸ	=	½ mv²
Ep	=	mgh
Eh	=	cm∆T
Eh	=	ml
Е	=	Pt

All forms of energy can be converted into any other form, so each of these equations can be equated to any other.

Example:

A skier of mass 60 kg slides from rest down a slope of length 20 m. The initial height of the skier was 10 m above the bottom and the final speed of the skier at the bottom of the ramp was 13 ms⁻¹.



Calculate:

- (a) the work done against friction as the skier slides down the slope;
- (b) the average force of friction acting on the skier.

Solution

(a)	Calculate E_p= mgh to work out	(b) Use Ew = Fd
	the amount of energy converted.	5880 = F x 20
		F = 5880/20
	E _p = mgh	F = 294N
	E _p = 60 x 9.8 x 10	
	E _p = 5880 J	

1.3 Collisions, Explosions and Impulse

Momentum

Conservation of Momentum

Momentum is the measure of an object's motion and is the product of mass and velocity.

p=mv

Since velocity is a vector so is momentum, therefore momentum requires a direction and we must apply the convention of + ve and – ve directions (typically we take direction to the right as positive).

An object can have a large momentum for two reasons, a large mass or a large velocity.

The law of conservation of linear momentum can be applied to the interaction (collision) of two objects moving in one dimension:

In the absence of net external forces, the total momentum before a collision = the total momentum after a collision

Collisions

The law of conservation of momentum can be used to analyse the motion of objects before and after a collision and an explosion. Let's deal with collisions first of all.

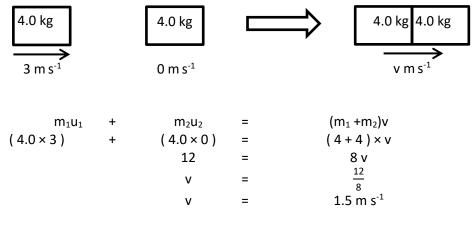
A collision is an event when two objects apply a force to each other for a relatively short time.

Example:

A trolley of mass 4.0 kg is travelling with a speed of 3 m s⁻¹. The trolley collides with a stationary trolley of equal mass and they move off together.

Calculate the velocity of the trolleys immediately after the collision.

Solution

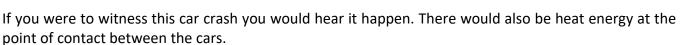


Kinetic Energy – Elastic and Inelastic Collisions

Elastic and Inelastic Collisions

When two objects collide their momentum is **always** conserved but, depending on the type of collision, their kinetic energy may or may not be. Take the two examples below:





These two forms of energy will have come from the kinetic energy of the cars, converted during the collision.

Here, kinetic energy is not conserved as it is lost to sound and heat. This is an inelastic collision.

In an elastic collision E_k before ≠ E_k after

2.

 $\rightarrow \leftarrow$ е е

When these two electrons collide they will not actually come into contact with each other, as their electrostatic repulsion will keep them apart while they interact.

There is no mechanism here to convert their kinetic energy into another form and so it is **conserved** throughout the collision.

In an elastic collision E_k before = E_k after

Elastic and Inelastic Collisions (continued)

Example:

A car of mass 2000 kg is travelling at 15 m s⁻¹. Another car, of mass 1500kg and travelling at 25 m s⁻¹ collides with it head on. They lock together on impact and move off together.

- (a) Determine the speed and direction of the cars after the impact.
- (b) Is the collision elastic or inelastic? Justify your answer.

Solutio	n 2000 kg		1500 kg € 25 m s ⁻¹	[2000 kg 1500 kg v m s ⁻¹
	m₁u₁ (2000 × 15) 30,000		m2u2 (1500 × (-25)) (- 37,500) v	= = = =	(m ₁ + m ₂)v (2000 + 1500) × v 3500 v -7500 / 3500 - 2.1 m s ⁻¹
The ne	gative sign mean	is the	cars are travelling at	t 2.1 m	s ⁻¹ to the left.
(b)					
	E _k before	=	½m₁u²		$+$ $\frac{1}{2}m_2u^2$
		=	-		+ $(\frac{1}{2} \times 1500 \times 25^2)$
		=	2.25 × 10 ⁵	4	+ 4.69 × 10 ⁵
		=	6.94 × 10 ⁵ J		
	E _k after	=	½m _{tot} v ²		
		=	½ × 3500 × 2.1 ²		
		=	7.72x10 ³ J		
	Kinetic ene	rgy ha	as not been conserve	ed, the	refore the collision is inelastic.

Explosions

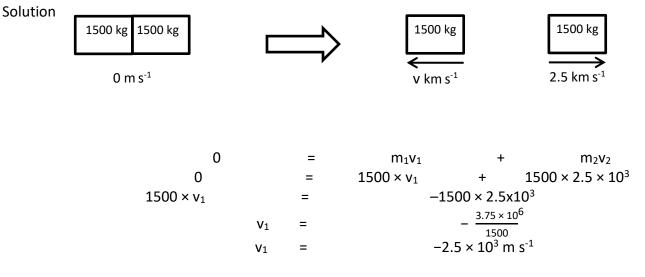
In a simple explosion two objects start together at rest then move off in opposite directions. Momentum must still be conserved, as the total momentum before is zero, the total momentum after must also be zero.

Example

An early Stark Jericho missile is launched vertically and when it reaches its maximum height it explodes into two individual warheads.

Both warheads have a mass of 1500 kg and one moves off horizontally, with a velocity of 2.5 km s⁻¹ (Mach 9) at a bearing of 090°.

Calculate the velocity of the other warhead.



The negative sign in the answer indicates the direction of v_1 is opposite to that of v_2 , i.e. 270° rather than 090°.

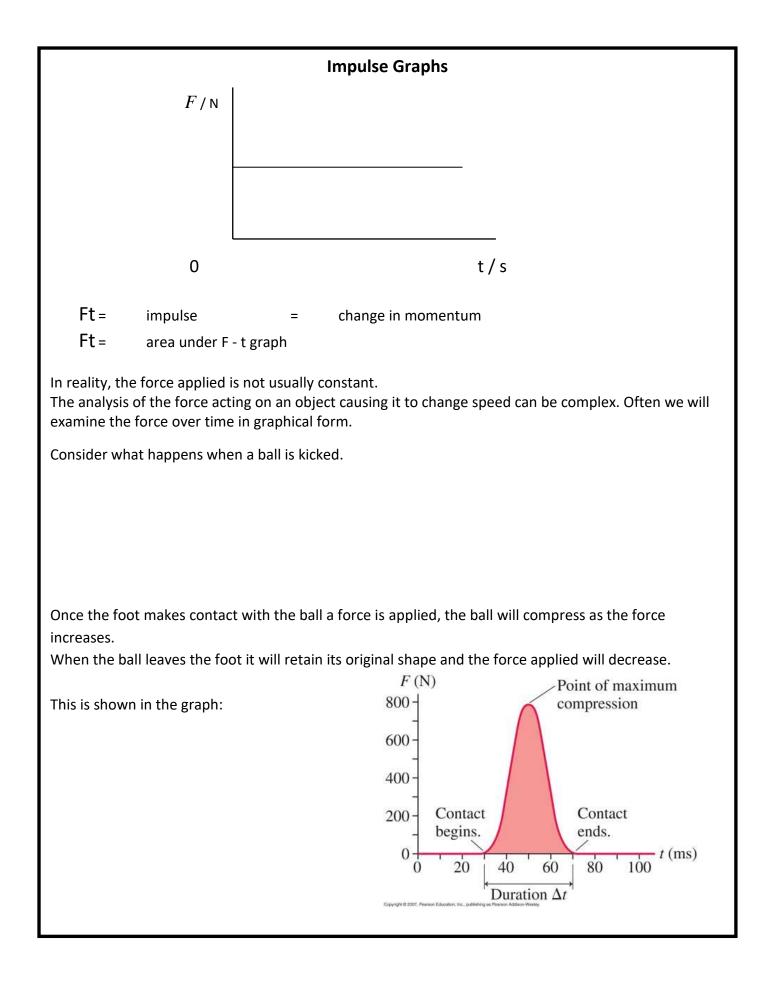
Second warhead is travelling at 2.5 km s⁻¹ on a bearing of 270° .

Newton's Third Law and Momentum		
Collisions	Collisions	
p_{total} before = p_{total} after	p_{total} before = p_{total} after	
$m_1u_1 + m_2u_2 = m_1v_1 + m_2v_2$	$0 = m_1 v_1 + m_2 v_2$	
$m_1u_1 - m_1v_1 = m_2v_2 - m_2u_2$	$m_1v_1 = -m_2v_2$	
$-m_1(v_1-u_1) = m_2(v_2-u_2)$	$\frac{m_1v_1}{t} = \frac{-m_2v_2}{t}$	
$-\underline{m_1(v_1 - u_1)} = \underline{m_2(v_2 - u_2)}$	$m_1a_1 = -m_2a_2$	
t t	$F_1 = -F_2$	
$-m_1a_1 = m_2a_2$		
$-F_1 = F_2$		

Impulse

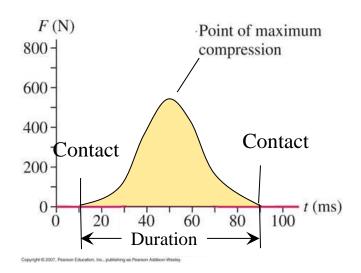
Impulse

Impulse is the product of	force and time, measured in	N s. Impulse is the cause of a change in momentum.		
	Impulse = Ft			
Impulse is equal to the c	nange in momentum, measu	red in kg m s ⁻¹ .		
	Impulse = change in mo Ft = Δp Ft = mv – mu (i	omentum mpulse has no symbol of its own)		
This means you can calcu	llate the impulse from:	F×t (Ns), or mv−mu (kg ms ⁻¹).		
A change in momentum	depends on:			
The size of the foThe time the force				
Beware – of objects which	h change directions!! 1 velo	city must be positive and 1 must be negative!!		
Example: A force of 100 N is applied to a ball of mass 150 g for a time of 0.020 s. Calculate the final velocity of the ball.				
Solution				
Solution:				
$F = 100N m = 0.150 kg t = 0.020 s u = 0 ms^{-1}$				
Ft = 100 × 0.020 = 2.0 = v = v =	mv – mu 0.150 × (v – 0) 0.150 v 2/0.15 13.3 ms ⁻¹			



Impulse Graphs (continued)

If a ball of the same mass that is softer is kicked and moves of with the same speed as that above, then a graph such as the one below will be produced.



The maximum force applied is smaller but the time it is applied has increased. Since the ball has the same mass and moves off with the same speed its impulse will be the same as the original.

This results in a graph of the same area but different configuration.

Example

A tennis ball of mass 100 g, initially at rest, is hit by a racquet.

The racquet is in contact with the ball for 20 ms and the force of contact varies over this period, as shown in the graph.

Determine the speed of the ball as it leaves the racquet.

Solution

Impulse = area under graph $= \frac{1}{2} \times 20 \times 10^{-3} \times 400$ = 4 N su = 0 m = 100 g = 0.1 kg v = ? 4 = 0.1v - (0.1 × 0) v = 40 ms^{-1}

Practical Applications – Car Safety

Essentially the greater the time you can take to decelerate an object, the smaller the force you need to apply. If your face is slowed by the dashboard the time to stop after you make contact with the dashboard will be small, resulting in a large force and a big OWWWWWW!!

Airbags

The concept of the airbag - a soft pillow to land against in a crash - has been around for many years. The first patent on an inflatable crash-landing device for airplanes was filed during World War II. In the 1980s, the first commercial airbags appeared in cars. Stopping an object's momentum requires a force acting over a period of time. When a car crashes, the force required to stop an object is very large because the car's momentum has changed instantly while the passengers' has not, there is not much time to work with. The goal of any restraint system is to help stop the passenger while doing as little damage to him or her as possible. What an airbag wants to do is to slow the passenger's speed to zero with little or no damage. To do this it needs to **increase** the time over which the change in speed happens.

Crumple Zones

Placed at the front and the rear of the car, they absorb the crash energy developed during an impact. This is achieved by deformation. While certain parts of the car are designed to allow deformations, the passenger cabin is strengthened by using high-strength steel and more beams. Crumple zones delay the collision. Instead of having two rigid bodies instantaneously colliding, crumple zones **increase the time** before the vehicle comes to a halt. This **reduces the force** experienced by the driver and occupants on impact. The change in momentum is the same with or without a crumple zone.

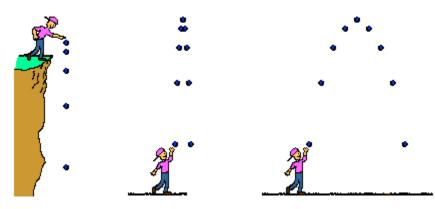


1.4 Gravitation

Projectile Motion

Projectiles

A projectile is any object, which, once projected, continues its motion by its own inertia and is influenced only by the downward force of gravity.



Most projectiles have both **horizontal and vertical** components of motion. As there is only a single force, gravity, acting in a single direction, which means only one of the components is being acted upon by the force. The two components are not undergoing the same kind of motion and must be treated **separately**.

Free Fall

air resistance

weight

When objects travel through the air, they have more than one force acting acting on them.

When an object is allowed to fall towards the Earth it will accelerate because of the force acting on it due to gravity, its **weight**. This will not be the only force acting on it though. There will be an upwards force due to **air resistance**. Air resistance increases with speed; you may notice this if you increase your speed when cycling.

If an object is allowed to fall through a large enough distance then the air resistance force may increase to become the same magnitude as weight of the object. The forces are now balanced and the object will fall with constant velocity, known as **terminal velocity**.

Motion During Free Fall and Terminal Velocity

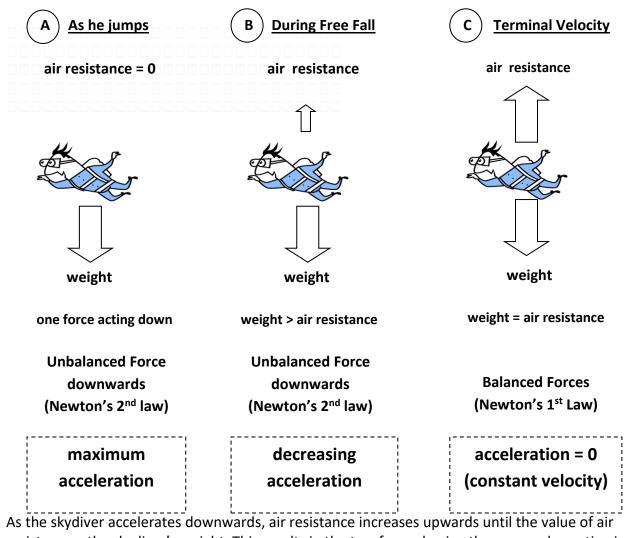
Free fall is the term used when an object is being acted upon only by the force of gravity (weight).

However, free fall will cause the object to accelerate and from the section **Introduction to Friction**, it states that:

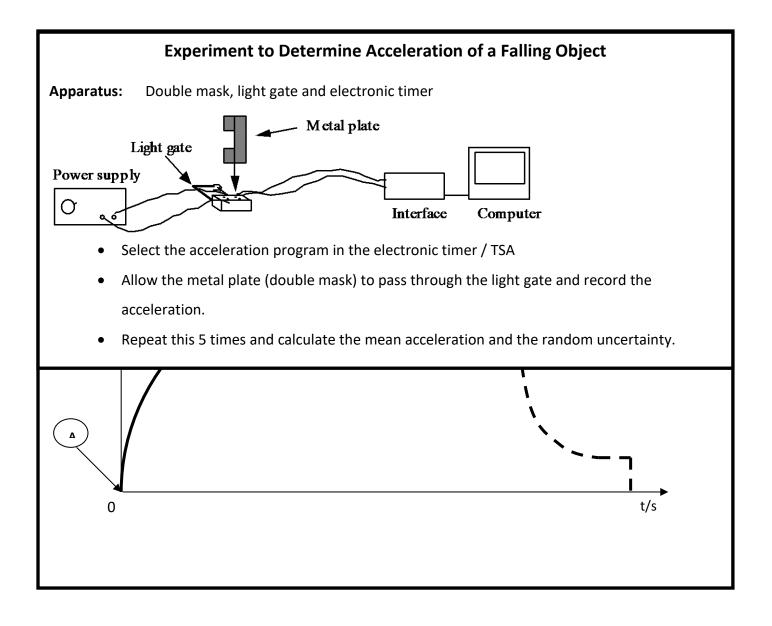
air resistance will increase as an object accelerates

so we have to look at air resistance as well as weight to study motion during free fall.

Example: If we consider a sky diver jumping out of an aeroplane



As the skydiver accelerates downwards, air resistance increases upwards until the value of air resistance = the skydiver's weight. This results in the two forces having the same value acting in opposite directions. According to Newton's 1st Law of Motion the skydiver will now travel at a constant velocity. This velocity is known as **Terminal Velocity**.



Horizontal Projection

Here is a classic horizontal projectile scenario, from the time of Newton.

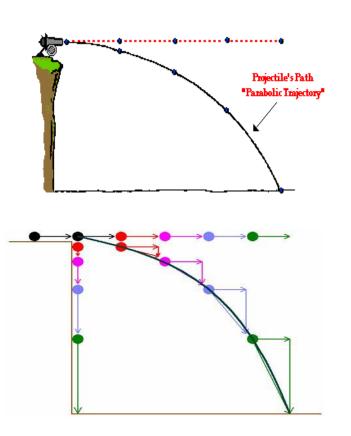
In projectile motion we **ignore all** air resistance, or any force other than gravity.

Analysis of this projectile shows the two different components of motion – these are both independent of each other.

Horizontally: there are no forces acting on the cannonball and therefore the horizontal velocity is constant.

Vertically: The force due to gravity is constant in the vertical plane and so the cannonball undergoes **constant vertical acceleration**.

The combination of these two motions causes the **curved** path of a projectile.



Example

The cannonball is projected horizontally from the cliff with a velocity of 100 m s⁻¹. The cliff is 20 m high. Determine:

- (a) the vertical speed of the cannonball, just before it hits the water;
- (b) if the cannonball will hit a ship that is 200 m from the base of the cliff.

Solution (Hint: time is the only quantity that	t can cross the horizontal and vertical barrier. Calculate t on one
side and use it on the other)	

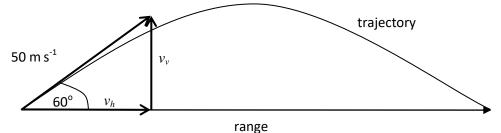
	Horizontal (use d = vt)				Vertical (use 3 equations of motion)				
	d = ? m				u = 0 ms ⁻¹				
v = 100 ms ⁻¹				v = ? ms ⁻¹					
				$a = 9.8 \text{ ms}^{-2}$					
				s = 20 m					
	t = ? s	5			t = ? s				
		= = =	vt 100 × 2∙02 ◀∽ 202 m vill hit the		Use t = 2.02 s on the horizontal	(a) (b)	19∙8 t	= = = = =	$0^{2} + 2 \times 9.8 \times 20$ 392 19.8 ms ⁻¹ u + at 0 + (9.8 × t) 19.8 / 9.8
ship.				side		t	=	2·02 s	

Projection at an Angle

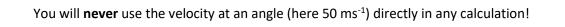
Projectiles at an angle are an application for our knowledge of splitting vectors into their horizontal and vertical components.



Example: The athlete has thrown the javelin at a velocity of 50 ms⁻¹ at an angle of 60° to the horizontal.



There is still only the single force of gravity acting on the projectile, so horizontal and vertical motions must still be treated very separately. This means that the velocity at an angle **must** be split into its **vertical and horizontal** components before any further consideration of the projectile.



Points to remember!!!

For projectiles fired at an angle above a horizontal surface:

1. The path of the projectile is symmetrical, in the horizontal plane, about the highest point. This means that:

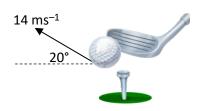
initial vertical velocity = – final vertical velocity
$$u_{\nu} = -v_{\nu}$$

- 2. The time of flight = $2 \times$ the time to highest point.
- 3. The vertical velocity at the highest point is zero.

Projection at an Angle Calculation

Example

A golfer hits a stationary ball and it leaves his club with a velocity of 14 ms^{-1} at an angle of 20° above the horizontal.



- (a) Calculate:
 - (i) the horizontal component of the velocity of the ball;
 - (ii) the vertical component of the velocity of the ball.
- (b) Calculate the time for the ball to reach its maximum height.
- (c) Calculate the total time of flight of the ball

v_H = 13.1 ms⁻¹

(d) How far down the fairway does the ball land?

Solution

	Horizontal	Vertical
	d = ? m	$u_v = 4.8 \text{ ms}^{-1}$ $v = 0 \text{ ms}^{-1}$ (at top) $a = -9.8 \text{ ms}^{-2}$ s = ? m t = ? s
	v _H = 13·1 ms ^{−1}	$v = 0 \text{ ms}^{-1}$ (at top)
	t = ? s	a = -9·8 ms ⁻²
		s = ? m
		t = ? s
(а) (i) v _H = v cos Ө	(a) (ii) v _H = v sin Ө
	v _H = 14 cos 20	v _H = 14 sin 20

	v _H = 4.8	ms⁻¹	
(b)	v	=	ι

(b)	v	=	u + at
	0	=	$4.8 + (-9.8 \times t)$
	9∙8 × t	=	4.8
	t	=	4-9 9-9
	t	=	0∙49 s

(c) total time = 2 x 0.49 s = 0.96 s

Newton's Thought Experiment – Satellite's orbit as an Application of Projectiles

Isaac Newton, as well as giving us the three laws, came up with an ingenious thought experiment for satellite motion that predated the first artificial satellite by over 300 years.

Essentially Newton suggested that if a cannon fired a cannonball it would fall towards the Earth. If it was fired at ever higher speeds then at some speed it would fall towards the Earth but never land since the curvature of the Earth would be the same as the flight path of the cannonball. This would then be a satellite. You would need a high mountain and an enormous cannon, but it would work. A satellite is in free fall around a planet.

Gravity

There is much confusion in Physics of the difference between mass and weight. This could be because the word weight is used, wrongly, in everyday life. People are always talking about their 'weight'. The need to 'lose weight' or 'gain weight'.

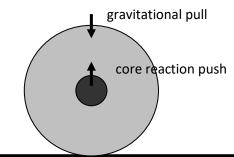
What they are really talking about is their *mass*. This is a measure of how much matter an object contains. This will only change if matter is added to or taken from the object. We can see this on a person as their body shape changes as matter is added or taken away.



What is Gravity?

Gravity is caused by mass. Any object that has mass will have its own gravitational field. The magnitude of the field depends on the mass of the object. Now, you are thinking that this is nonsense. If this were the case then everything on Earth would have it's own gravity and that's just not true. Well, actually, it is true. It's just that those gravitational forces are so small that we don't notice them. Remember, the gravitational pull of the Earth is 9.8 Nkg⁻¹ and the Earth has a mass of 5.97 × 10^{24} kg just think how tiny the gravitational pull you exert must be.

Gravity is a force that permeates the entire universe; scientists believe that stars were formed by the gravitational attraction between hydrogen molecules in space. The attraction built up, over time, a large enough mass of gas such that the forces at the centre of the mass were big enough to cause the hydrogen molecules to fuse together, generating energy. This is what is happening in the centre of the sun. The energy radiating outwards from the centre of the sun counteracts the gravitational force trying to compress the sun inwards.



In time the hydrogen will be used up, the reaction will stop and the sun will collapse under its own gravity. If you expect to live for 4 or so billion years you could worry about this.

Newton's Universal Law of Gravitation

Newton produced what is known as the Universal Law of Gravitation

$$F = \frac{Gm_1m_2}{r^2}$$

G is the universal constant of gravitation = $6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$.

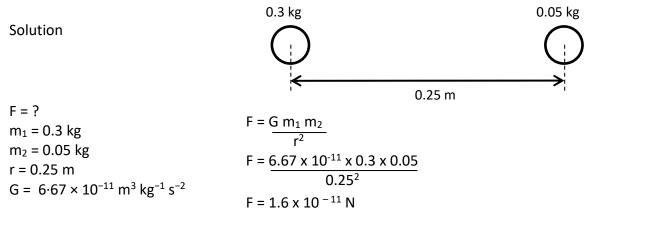
Newton's Law of gravitation states that the gravitational attraction between two objects (m_1 and m_2) is directly proportional to the mass of each object and is inversely proportional to the square of their distance (r) apart.

Gravitational force is always attractive, unlike electrostatic or magnetic forces.

The distance r between the two objects is the distance between their **centres of mass**. This is especially important when considering planetary bodies. For example, the radius of the orbit of the moon is only the distance from the surface of the Earth to the surface of the Moon, not the distance between their centres of mass.

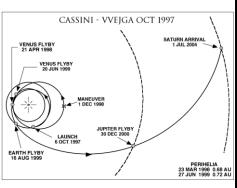
Example

Consider a folder, of mass 0.3 kg and a pen, of mass 0.05 kg, sitting on a desk, 0.25 m apart. Calculate the magnitude of the gravitational force between the two masses. (Assume they can be approximated to spherical objects).



Application of Gravitational Force – The 'Slingshot Effect'

Another application of the gravitational force is the use made of the 'slingshot effect' by space agencies to get some 'free' energy to accelerate their spacecraft. Simply put they send the craft close to a planet, where it accelerates due to the gravitational field of the planet. Here's the clever part, if the trajectory is correct the craft then speeds past the planet with the increased speed. Don't get it right and you still get a spectacular crash into the planet, could be fun but a bit on the expensive side!



1.5 Special Relativity Relativity

Reference Frames

Relativity is all about observing events and measuring physical quantities, such as distance and time, from different reference frames. Here is an example of the same event seen by three different observers, each in their own frame of reference:

Event 1: You are reading your Kindle on the train. The train is travelling at 60 mph.

Observer	Location	Observation
1	Passenger sitting next to you	You are stationary
2	Person standing on the platform	You are travelling towards them at 60 mph
3	Passenger on train travelling at 60 mph in opposite direction	You are travelling towards them at 120 mph.

This example works well as it only involves objects travelling at relatively low speeds. The comparison between reference frames does not work in quite the same way, however, if objects are moving close to the speed of light.

Event 2: You are reading your Kindle on an interstellar train. The train is travelling at 2×10^8 ms⁻¹.

Observer	Location	Observation
1	Passenger sitting next to you	You are stationary
2	Person standing on the platform	You are travelling towards them at 2 × 10 ⁸ ms ⁻¹
3	Passenger on train travelling at 2 × 10 ⁸ ms ⁻¹ in opposite direction	You are travelling towards them at 4 × 10⁸ ms ⁻¹

The observation made by observer 3 is impossible as an object cannot travel faster than the speed of light in any reference frame and it would certainly be impossible to watch something travel faster than light, so this scenario is impossible.

The Principles of Relativity - Introduction

Using his imagination and performing thought experiments (gedanken) like those above, Einstein came up with two principles, or postulates, to explain the problem of fast moving reference frames. These were later proved with a vast array of data from many different experiments and became very clear once we started communicating with satellites, in orbit.

The postulates of Special Relativity:

- 1. When two observers are moving at constant speeds relative to one another, they will observe the same laws of physics.
- 2. The speed of light (in a vacuum) is the same for all observers.

This means that no matter how fast you go, you can never catch up with a beam of light, since it always travels at 3.0×10^8 ms⁻¹ relative to you.

If you (or anything made of matter) were able to travel as fast as light, light would still move away or towards you at 3.0×10^8 ms⁻¹, as you are stationary in your own reference frame.

The most well-known experimental proof is the Michelson-Morley interferometer experiment. Maxwell's electromagnetism equations also corroborated these postulates.

Example: If a car ship is travelling through space at 90% of the speed of light and then switches on its

headlights. The passenger of the car will see the beams of the headlights travel away from them at 3×10^8 ms⁻¹. An observer on Earth will also observe light of the beams travelling at 3×10^8 ms⁻¹.

The speed of light, **c**, is constant in and between all reference frames and for all observers.

These principles have strange consequences for the measurement of distance and time between reference frames.



Time Dilation

Time Dilation Equation and the Lorentz Factor

$$t^{'}=\frac{t}{\sqrt{1-\left(\frac{v^{2}}{c^{2}}\right)}}$$

t' is always observed by the <u>stationary</u> observer, observing the object moving at speed. E.g. the person on a train platform watching the train go by, or an observer on Earth watching a fast moving ship.

Note this is often written as: $t' = t\gamma$

where \mathbf{y} is known as the Lorentz Factor. It is used often in the study of special relativity and is given by:

$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v^2}{c^2}\right)}}$$

Example:

A rocket is travelling past Earth at a constant speed of 2.7×10^8 ms⁻¹. The pilot measures the journey as taking 240 minutes. How long did the journey take when measured by an observer on Earth?

Solution:

t = 240 minutes c = $3 \times 10^8 \text{ ms}^{-1}$ v = $2.7 \times 10^8 \text{ ms}^{-1}$ t' = ?

$$t' = \frac{240}{\sqrt{1 - \left(\frac{2.7 \times 10^8}{3.0 \times 10^8}\right)^2}}$$

 $t' = \frac{t}{\sqrt{1 - \left(\frac{v^2}{2}\right)}}$

t' = 550 minutes

An observer on Earth would measure the journey as taking 550 minutes.

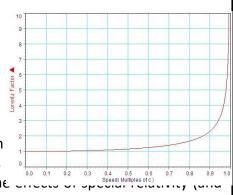
Why we do not notice relativistic time differences in everyday life?

A graph of the Lorentz factor versus speed (measured as a multiple of the speed of light) is shown below.

We can see that for small speeds (i.e. less than 0.1 times the speed of light) the Lorentz factor is approximately 1 and relativistic effects are negligibly small. Even 0.1 times the speed of light is 300,000 ms⁻¹ or 1,080,000 km h⁻¹ or about 675,000 mph – a tremendously fast speed compared to everyday life.

However, the speed of satellites is fast enough that even these small changes will add up over time and seriously affect the synchronisation of global positioning systems (GPS) and television satellites with users

on the Earth. They have to be specially programmed to adapt for the critical of these small changes in also general relativity, which is not covered here). Very precise measurements of these small changes in time have been noted on fast-flying aircraft and agree with predicted results within experimental error.



Application of Time Dilation

Further evidence in support of special relativity comes from the field of particle physics, in the form of the detection of a particle called a muon at the surface of the Earth. Muons are produced in the upper layers of the atmosphere by cosmic rays (high-energy protons from space). The speed of muons high in the atmosphere is 99.9653% of the speed of light.

The half-life of muons when measured in a laboratory is about $2 \cdot 2 \ \mu s$.

Example: Show, by calculation, why time dilation is necessary to explain the observation of muons at the surface of the Earth.

Solution:

t = $2.2 \ \mu s = 2.2 \times 10^{-6} s$ v = $0.999653 \times 3.00 \times 10^{8} = 2.998956 \times 10^{8} ms^{-1}$ d = ?

d = vt d = $2.998956 \times 10^8 \times 2.2 \times 10^{-6}$ d = 660 m

$$t' = \frac{t}{\sqrt{1 - \left(\frac{v^2}{c^2}\right)}}$$
$$2.2 \times 10^{-6}$$

$$t = \frac{1}{\sqrt{1 - (0.999653)^2}}$$

t' = 84 s
d' = vt
d' = 2.998956 ×
$$10^8 \times 84 \times 10^{-6}$$

d' = 2.52 × 10^4 m

In the reference frame of an observer on Earth the half-life of the muon is recorded as $84 \mu s$ and therefore from this perspective, the muon has enough time to travel the many kilometres to the Earth's surface.

A Twin Paradox

You leave Earth and your twin to go on a mission in a spaceship travelling at 90% the speed of light on a return journey that lasts 20 years. When you get back you find that 46 years will have elapsed on Earth. Your clock will have run slowly compared to one on Earth, however as far as you were concerned the clock would have been working correctly on your spaceship.

You will look 26 years younger than your twin.

Length Contraction

Another implication of Einstein's theory is the shortening of length when an object is moving. Consider the muons discussed above. Their large speed means they experience a longer half-life due to time dilation. An equivalent way of thinking about this is that the fast moving muons observe a much shorter (or contracted) distance travelled, by the same amount as the time has increased (or dilated). A symmetrical formula for length contraction can be derived.

$$I' = I \sqrt{1 - \frac{v^2}{c^2}}$$

Where I is the distance measured by an observer who is stationary and I' the distance observed by the observer who is moving at speed.

<u>Example</u>

Let's take the example of a space ship flying away from Earth towards Proxima Centauri, our nearest star, to study the observations due to length contraction. The distance to Proxima Centauri is 4.2 ly. Length contraction only takes place in the direction that the object is travelling. For the pilot of the space ship, this means that they will measure the distance, in front of them, between Earth and Proxima as less than the distance measured by a stationary observer.

 v^2

Let's say the spaceship is travelling at 0.8 c.

l = 4.2 ly l' = ?	$I' = I \sqrt{1 - \frac{v}{c^2}}$
V = 2 v = 0.8 c	$i' = 4.2 \sqrt{1-0.8^2}$
	l ['] = 2.52
	l [′] = 2.5 ly

So the Pilot of the ship measures their journey as 2.5 ly.

1.6 The Expanding Universe **The Doppler Effect and Redshift**

The Doppler Effect

The Doppler Effect is the change in the observed frequency of a wave (both light and sound), when the source or observer is moving.

In this course we will concentrate on a wave source moving at constant speed relative to a stationary observer.

You have already experienced the Doppler Effect many times. The most noticeable is when a police car, ambulance or fire engine passes you. You hear *i* the pitch of their siren increase as they come towards you and then decrease as they move away. Another memorable example is the sound of a very fast moving vehicle, such as a Formula 1 car passing you (or passing a microphone on the television), the sound of the engine rises and falls in frequency as it approaches, passes and moves away.

The Doppler Effect applies to all waves, including light.

Uses of the Doppler Effect

- Police radar guns use the Doppler effect to measure the speed of motorists.
- Doppler is used to measure the speed of blood flow in veins to check for deep vein thrombosis [DVT] in medicine.

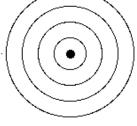
Stationary Source

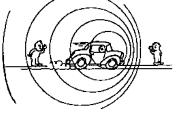
A stationary sound source produces sound waves at a constant frequency f, and the wavefronts propagate symmetrically away from the source at a constant speed, which is the speed of sound in the medium. The distance between wave-fronts is the wavelength. All observers will hear the same frequency, which will be equal to the actual frequency of the source: $f = f_0$.

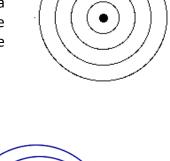
Moving Source

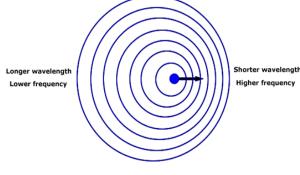
The sound source now moves to the right with a speed \mathbf{v}_{s} . The wavefronts are produced with the same frequency as before, therefore the period of each wave is the same as before. However, in the time taken for the production of each new wave the source has moved some distance to the right. This means that the wavefronts on the left are created further apart and the wavefronts on the right are created closer together. This leads to the spreading out and bunching up of waves you can see to the right and hence the change in frequency.

The frequency of the source will remain constant, it is the observed frequency that changes.









The Doppler Effect Equations

More relevant to our learning in this section, the Doppler Effect is highly prominent in our observations of the universe and provides some of the strongest evidence for major theories such as the Big Bang and an expanding universe.

For a stationary observer with a wave source **moving towards them**, the relationship between the frequency, \mathbf{f}_s , of the source and the observed frequency, \mathbf{f}_o , is:

$$f_{o} = f_{s} \left(\frac{v}{v - v_{s}}\right)$$

$$v = speed of the wave$$

$$v_{s} = speed of source$$

$$f_{s} = frequency source$$

$$f_{o} = observed frequency$$

For a stationary observer with a wave source **moving away from them**, the relationship between the frequency, f_s , of the source and the observed frequency, f_o , is:

$$f_o = f_s \left(\frac{v}{v - v_s}\right)$$

This second scenario is exactly what is observed when we look at the light from distant stars, galaxies and supernovae, evidence that the universe is expanding. These relationships also allow us to calculate the speed at which an exoplanet is orbiting its parent star, or the velocity of stars orbiting a galactic core, which has lead us to theorise the existence of dark matter.

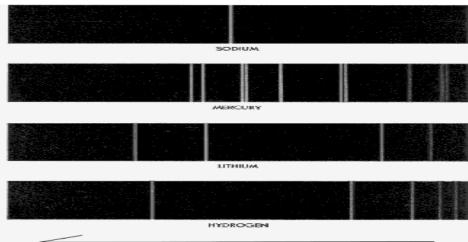
An Example of the Doppler Effect - Red Shift

Redshift is an example of the Doppler Effect. The light from stars, as observed from Earth, is always reduced in frequency and shifted towards the red (longer wavelengths) end of the spectrum. This is because the stars and galaxies are sources of light which are moving away from us.

Redshift has always been present in the light reaching us from stars and galaxies but it was first noticed by astronomer Edwin Hubble, in the 1920's, when he observed that the light from distant galaxies was shifted to the red end of the spectrum (longer wavelengths).

The light emitted by a star is made up of the line spectra emitted by the different elements present in that star. Each of these line spectra is an identifying signature for an element and these spectra are constant throughout the universe. You will learn a lot more about spectra in the Particles and Waves unit of this course.

Since these line spectra are so recognisable, we can compare the spectra produced by these elements, on Earth, with the spectra emitted by a distant star or galaxy.



Examples of line spectra of different elements

Wavelength

Hubble examined the spectral lines from various elements and found that the spectra emitted by each galaxy were shifted **towards the red** by a specific amount. This **shift** was due to the **galaxy moving away from the Earth** at speed, causing the Doppler Effect to be observed. The **bigger the magnitude of the shift the faster the galaxy** was moving.

Redshift of a Galaxy Equation

Redshift, z, of a galaxy is given by:

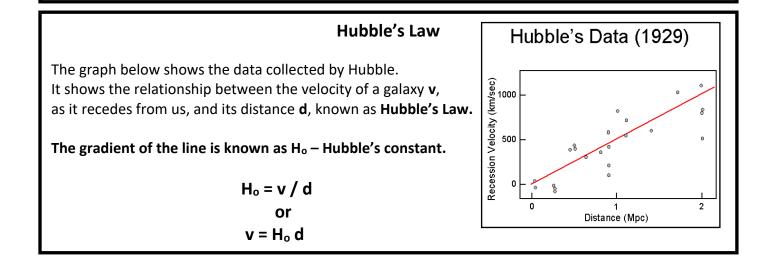
$$z = \frac{\lambda_{observed} - \lambda_{rest}}{\lambda_{rest}} = \frac{\Delta\lambda}{\lambda_{rest}}$$

Redshift of slow moving galaxies, can also be shown to be the ratio of the (recessional) velocity of the galaxy to the velocity of light:

$$z = \frac{v_{galaxy}}{c}$$

As redshift is always calculated from the ratio of quantities with the same unit, it has **no unit of its own**.

Over the course of a few years Hubble examined the red shift of galaxies at varying distances from the Earth. He found that the further away a galaxy was the faster it was travelling away from us. The relationship between distance and speed of a galaxy is known as Hubble's Law.



Calculating the Age of the Universe

Hubble's observations show that galaxies are moving away from the Earth and each other in all directions, which suggests that the universe is expanding. This means that in the past the galaxies were closer to each other than they are today. By working back in time it is possible to calculate a time when all the galaxies were at the same point in space. This allows the age of the universe to be calculated.

- v = speed of galaxy receding from us
- **d** = distance of galaxy from us
- H₀ = Hubble's constant
- t = time taken for galaxy to travel that distance, i.e. the age of the Universe

$$t = \frac{d}{v} \qquad (v = H_o d)$$

$$t = \frac{d}{H_o d}$$

$$t = \frac{1}{H_0}$$

Currently, using this method, NASA estimate the age of the universe to be **13·7 billion years**. Since Hubble's time, there have been other major breakthroughs in astronomy and our ability to make accurate observations of very distant objects. All of these support the findings of Hubble, but allow the age of the Universe to be calculated even more accurately.

Evidence for the Expanding Universe

It is generally accepted, based on the evidence given previously, that the Universe is expanding. The evidence for this includes:

Velocity of Galaxies

There are many more galaxies showing redshift (travelling away from us) than those exhibiting blueshift (i.e. travelling towards us).

Dark Matter

How can we measure the mass of objects in space? You would need a big set of scales. In fact astronomers can relate the orbital speed of galaxies to their masses. The problem is that the masses measured seem to be bigger than the mass that can be accounted for by the number of stars present in a galaxy. This leads to the theory of '**Dark Matter**'. Basically there appears to be stuff there that we can't see and don't know what it is, so for the moment give it a name and hope we find out what it actually is later.

Dark Energy

Now if that was the only thing it might not be too bad, but the universe is expanding at a greater rate than astronomers would expect. It seems that something appears to be opposing the gravitational force. This is called **Dark Energy**. This increased expansion appears to verify Einstein's inclusion of the cosmological constant in his Special Theory of Relativity.

Big Bang Theory

The universe started with a sudden appearance of energy which consequently became matter and is now everything around us.

The temperature of stellar objects

You will have seen what happens to a piece of iron as it is heated, as it gets hotter its colour changes from dull red to bright red to orange then yellow. This also happens with stellar objects. The temperature of stellar objects is related to the distribution of emitted radiation over wide range of wavelengths.

What this means is that by examining the spectrum of a distant star, its temperature can effectively be determined. This graph shows that

- The peak wavelength is shorter for hotter objects
- irradiance increases with temperature Hot objects emit more radiation per unit surface area per time than cooler objects.

Evidence for the Big Bang Theory

The Expanding Universe

As previously discussed, the expanding universe can be shown by calculations of redshirting galaxies and data collected by Hubble on recession velocity.

Cosmic Microwave Background Radiation

In 1989 a satellite was launched to study the background radiation, it was called the Cosmic Background Explorer [COBE]. In 1992 it was announced that COBE had managed to measure fluctuations in the background radiation. This was further evidence to support the Big Bang theory.

An image of the fluctuations is shown below.

The Abundance of Light Elements

Other evidence to support the Big Bang theory includes the relative abundances of hydrogen and helium in the Universe.

Scientists predicted that there should be a significantly greater proportion of hydrogen in the Universe. The next most abundant should be helium.

The elements present in the Universe can be determined by spectroscopy, which you will study later in unit 3.

The latest proportions are given in the table shown. These observations conform to the predicted proportions.

Evidence for the Big Bang Theory (continued)

The darkness of the Sky

Another is the explanation for Olber's paradox. His paradox was in answer to the question, "why is the sky dark at night?"

This is not as obvious as you first might imagine...

If the Universe followed the Steady State model then there should be an even distribution of stars in all directions. All the stars in the Universe should be visible. This means the light from the stars should reach Earth and the sky should be bright.

The Big Bang theory gives a finite age to the Universe, and only stars within the observable Universe can be seen. This means that only stars within the distance of 15,000 light years will be observed. Not all stars will be within that range and so the dark sky can be explained.